

# Sedimentation-driven coagulation inside the snow-line

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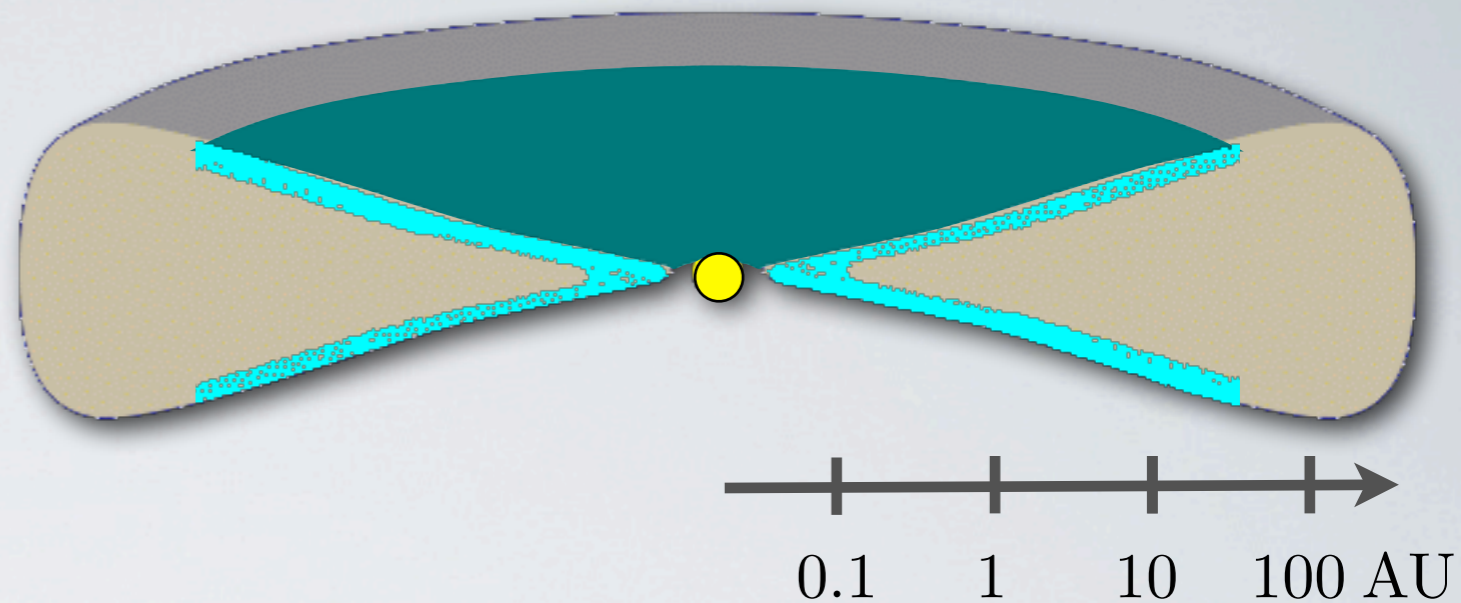
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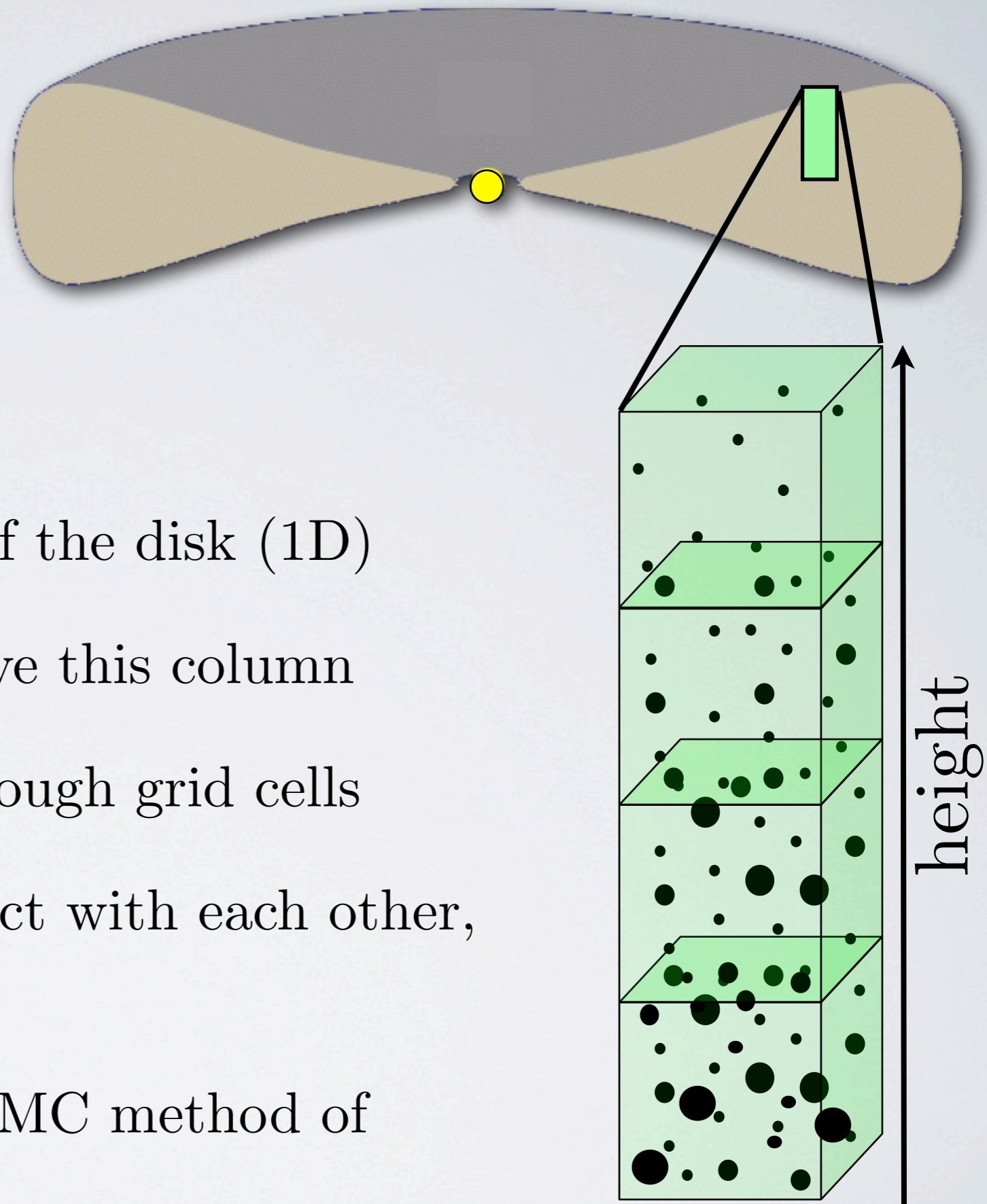
# Our goal



- micron sized particles present in disk atmospheres for  $\sim 10^6$  years (disk SED, 10 micron feature - Henning&Meeus, 2011)
- what prevents particle growth and settling to the midplane?
  - Brauer et al. 2008, Birnstiel et al. 2010: fragmentation produces micron sized particles, turbulence spreads the small particles vertically
  - Zsom et al. 2010: bouncing barrier prevents fragmentation, *the question remains unanswered?*
    - local simulations performed at the midplane of the disk

# Numerical set-up

- simulate a vertical column of the disk (1D)
- particles do not enter or leave this column
- particles move vertically through grid cells
- Monte Carlo particles interact with each other, if located in the same grid
- Particle growth followed by MC method of Zsom&Dullemond, 2008

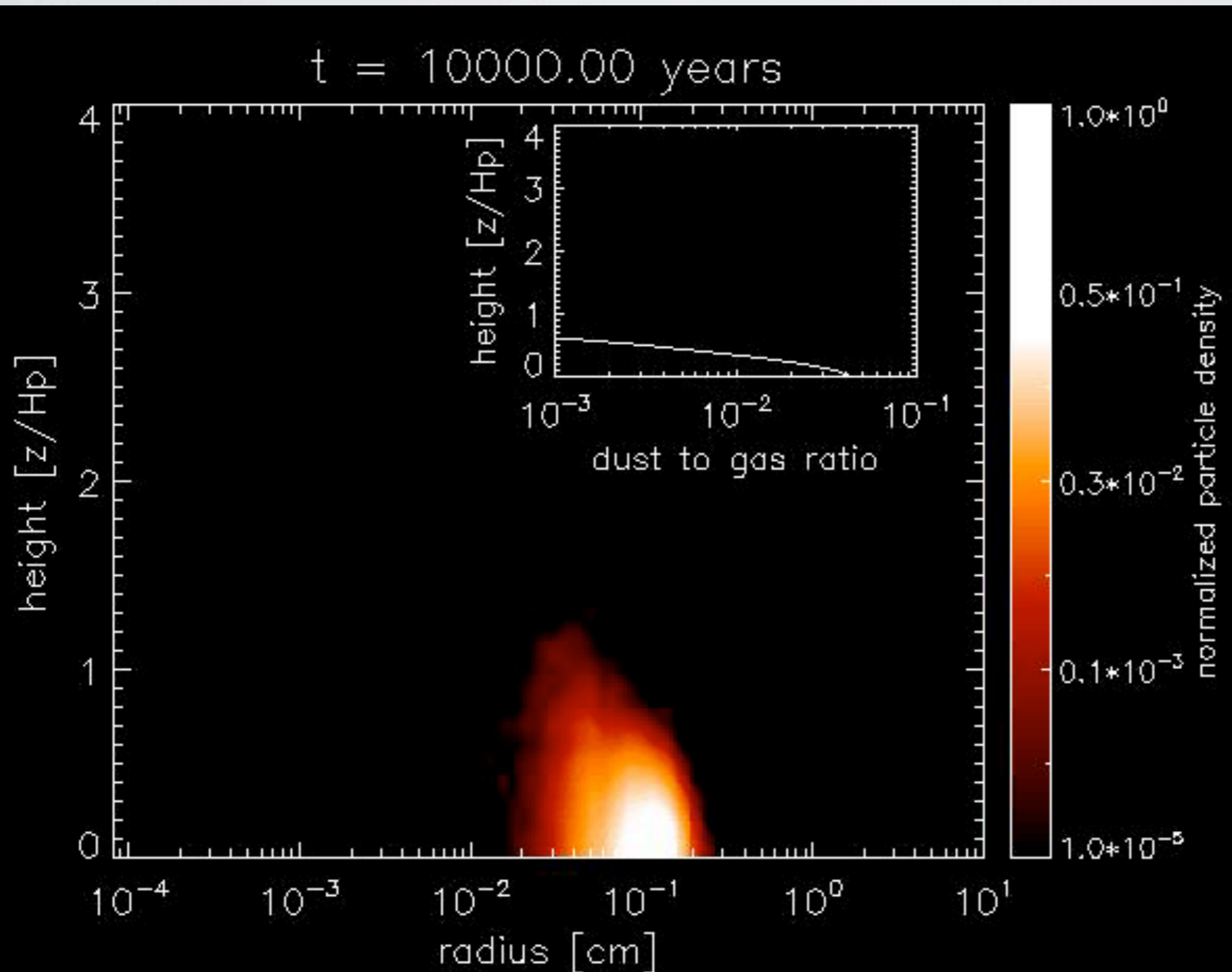


# Physical processes

- Particle motion and relative velocity sources:
  - Brownian motion, turbulence (Ormel&Cuzzi relative velocity, Shakura & Sunyaev alpha description), vertical settling
- Particle growth:
  - Collision model based on laboratory measurements (Güttler et al. 2010)
  - includes sticking, bouncing, fragmentation
- Particle structure:
  - hit&stick porosity model of Okuzumi et al 2009

# Results

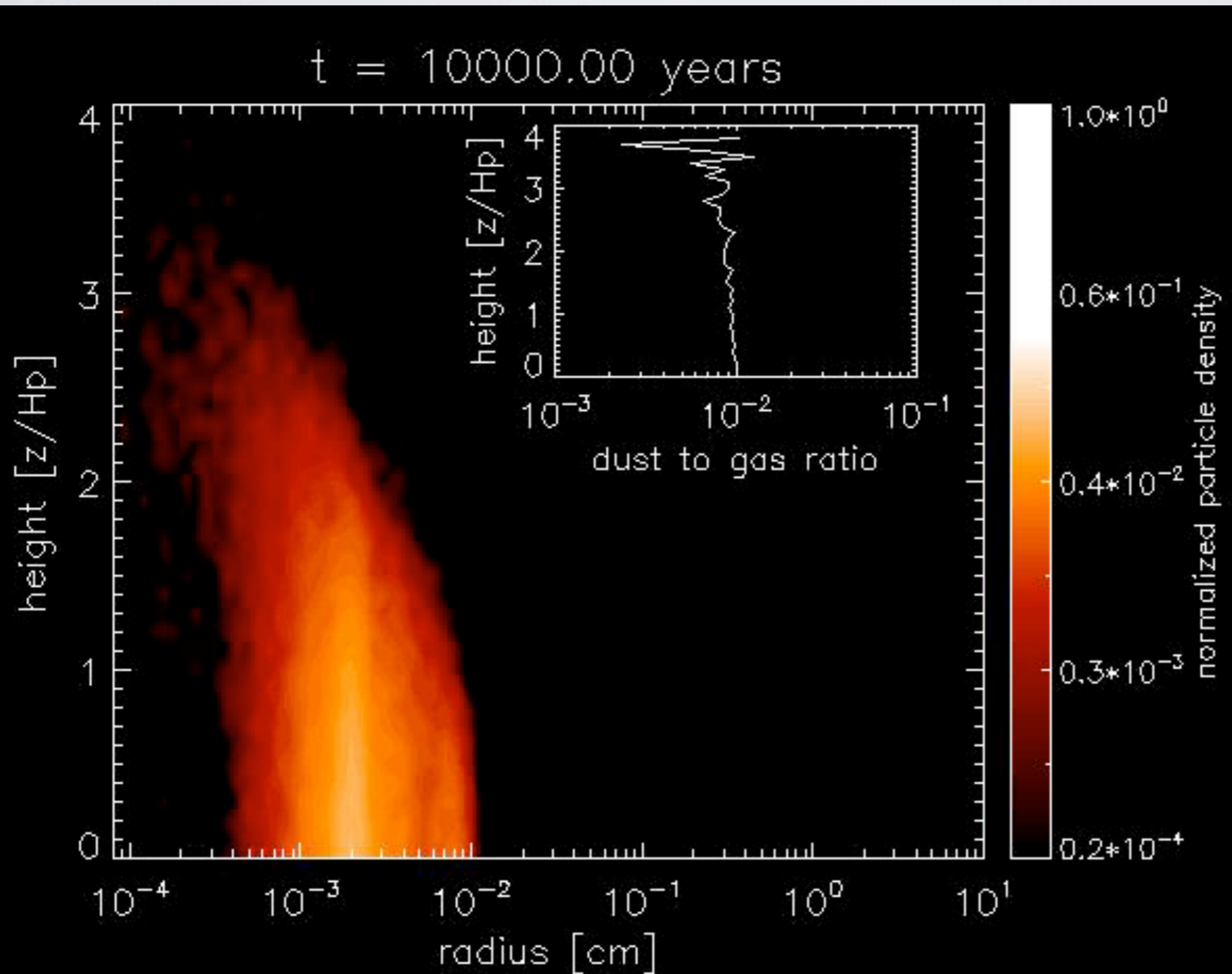
- Shakura & Sunyaev **alpha** turbulence parameter is  $10^{-4}$



$H_d = 0.2 H_g$   
size:  $10^{-1}$  cm

# Results

- Shakura & Sunyaev **alpha** turbulence parameter is  $10^{-2}$



$$H_d = 0.95 H_g$$

size:  $10^{-2}$  cm

# Conclusions

- Local simulations: to study the effects of unexplored processes
  - but does not necessarily indicate global behavior
- Density in local simulations of Zsom et al 2010: *constant*
  - does not produce small particles for any alpha
- Density in 1D simulation: *decreasing function of height*
- *This combined with high alpha produces small grains!*
- For more details see Zsom et al, 2011, 534, 73