Astrophysics Lab Work – Computational Basics

This exercise will help you to become acquainted with some basic features of Linux and gnuplot, in order to be able to solve the problems involving computational aspects in the various parts of your lab work.

The exercises can be performed either

i) recommended: on your own computer, by connecting with the numprakt accounts of some of the instute's machines (ltsp04 ... ltsp17) via tigervnc with active eduVPN tunnel.

NOTE: if your machine is "busy" (other users already logged in), try another machine. Try to work always on the same machine. If this is not possible, copy (via scp) your working directory (see Exercise 1) from your former machine to the currently used one.

- ii) on your own computer, if you have LINUX installed (either directly, or as a virtual system), and a working internet connection. To connect to our numprakt-machines, the eduVPN tunnel must be active.
- iii) on your own computer under any operation system, when you have an X-server running and can connect to the hosts according to i) via ssh. To connect to our numprakt-machines, the eduVPN tunnel must be active.

Passwords and machine/session numbers have been provided to you in the "intro"-lab.

Please log all your commands from exercise 1 and 2 (plus additional comments/information from the other exercises) to a file YourName_logfile.txt within your working directory (see below). Log only successful commands!

Exercise 1 – files and directories

Please log all commands (see above)

- Login to your home directory or to numprakt on the USM-hosts (see above)
- Create a directory called YourName, where YourName should be replaced by your actual name. This will be your working directory.
- In this directory, create a subdirectory Europe.
- Within this directory, create two subdirectories called Germany and Italy.
- Change to Germany and create the (empty) files Roma_I Freiburg_D Milano_I Salzburg_A Berlin_D Hamburg_D Ingolstadt_D and Wien_A.
- Move all files containing an I in their names to the directory Italy.
- Delete all files (within Germany) which contain a burg in their names.
- Rename file Wien_A to Stuttgart_D.

- Enter directory Italy and copy all files ending with _D to directory Germany.
- Delete all files which contain a _D.

Exercise 2 – file editing

For this exercise, use the editor emacs. Again, log all your commands (see above), and provide the commands used inside emacs.

Exercise 2a – line editing

- Open an ssh connection to numprakt@ltsp08.usm.uni-muenchen.de.
- List the content of public/Europe in long listing format onto your screen. Redirect this listing to a file called public/YourName_Europe.txt.
- Logout the ssh connection.
- Copy, via scp, the file public/YourName_Europe.txt from numprakt@ltsp08.usm.uni-muenchen.de to your working directory.
- Open this file with emacs and edit the content via *cut and paste* in such a way that all files with the ending .Germany appear first, followed by the files with ending .Italy. Delete all other lines, and save the file. For this task, do *not* use the commands from the menu bar, but learn to control emacs by keys. In the end, this works much faster.
- Convert the file to pdf (hint: first, create a ps file via a2ps).

Exercise 2b – column editing

- Again, log your commands, including those used inside emacs.
- Copy, via scp, the file public/wlr.dat on numprakt@ltsp08.usm.uni-muenchen.de to your working directory, under name YourName_wlr.dat, and open it with emacs.
- Delete columns 5 to 9, 11 to 15 and the last column (17), via the corresponding ..._rectangle command, and insert them below the modified table, allowing for one empty line in between.
- The new file should consist of two tables, the first one with 6 columns, the second one with 11 columns. The first table will be used in **Exercise 3b**, and contains the following information about a sample of Galactic O-stars:
 - 1. Name, identified by HD number (Henry-Draper catalogue),
 - 2. lc, luminosity class,
 - 3. Teff (effective temperature), in units of 1000 K,
 - 4. Rstar (stellar radius), in units of solar radius,
 - 5. vinf (terminal velocity of stellar wind), in units of km/s,
 - 6. Mdot (mass-loss rate of stellar wind), in units of 10^{-6} solar masses per year.

Rename the (modified) files from Exercise 2 to 1_Europe.pdf and 2_wlr.dat, respectively.

Exercise 3 – Gnuplot

In this exercise you will become acquainted with gnuplot. You needn't log the single commands. Again, use the editor emacs for editing. (If you are working in your own linux-environment, and you have not installed gnuplot, please install it)

Exercise 3a – demo tour and first edits

- locate the file-path to the directory gnuplot/demo on numprakt@ltsp08.usm.unimuenchen.de, denoted by \$GPATH in the following.
- Copy the directory **\$GPATH/gnuplot/demo** to a new one, **YourName_gnuplot**, in your working directory.
- Change to YourName_gnuplot and start gnuplot.
- Let gnuplot introduce itself, by starting the demo-tour with the command load 'all.dem'.
- Check for an example which impresses you most, locate the corresponding .dem file, and describe briefly the purpose of this script (including its name) in your log-file.
- Have a careful look into the capabilities of gnuplot. Move the mouse cursor over the plot, and type 'h'. Study the output and 'play around' with the figure. Consider, e.g., the possibilities to zoom in and out, replot the data, switch the grid on and off, switch the log scale(s) on and off etc. (<B1> to <B3> refer to the mouse buttons). When 3-D data are plotted, try to rotate the figures with the mouse.

For future work, you might use specific examples as templates for your own plots!

- Find out which .dem file creates the world map, and modify this file in such a way that only the 2-D world map is plotted (new filename = 3_world.dem). Create a corresponding postscript file (filename = 4_world.ps), and print the map.
- Modify the data in such a way that only "Munich" is located within the map, and create another ps-file (filename = 5_world_munich.ps) with this modification.

Exercise 3b – a first own plot

In this exercise, you will derive the so-called *wind-momentum luminosity relation* (WLR) for Galactic supergiants, which relates the momenta of their radiation driven winds, modified by the **sqrt** of their radii, with their luminosity. To this end, proceed as follows:

- Copy 2_wlr.dat to my_wlr.dat.
- In my_wlr.dat, delete the 2nd table. Within the first table, comment all entries with l.c.≠1 by inserting a hash # at the beginning of the corresponding lines. (The hash is the comment sign for shell scripts and gnuplot). In this way, only data for lc=1 objects (supergiants) will be processed by gnuplot.
- As a first test, plot radius as a function of Teff for the Galactic O-supergiants via gnuplot. Convince yourself that the correct number of objects are plotted. Save the ps-version of this plot under 6_ostars.ps (inspect it via gv).

In the following, you will calculate and plot the modified wind momenta of the stars as a function of their luminosity (log – log plot, data from file my_wlr.dat), and overplot a corresponding linear regression. The result will display the WLR for Galactic supergiants.

- Work with two windows in parallel: one for gnuplot, where you can test your solution, and one for emacs, where you create and modify a corresponding gnuplot-script, which should be named as my_wlr.gpl.
- By means of this script, plot the modified wind-momenta read and calculated from file my_wlr.dat (see Exercise 2b) as a function of luminosity as symbols, i.e., plot $\log_{10} D_{\text{mom}}$ as a function of $\log_{10} L/L_{\odot}$, where

$$D_{\text{mom}} = \dot{M} \cdot v_{\text{inf}} \sqrt{R_{\text{star}}/R_{\odot}}$$
$$L/L_{\odot} = (R_{\text{star}}/R_{\odot})^2 \cdot (T_{\text{eff}}/T_{\text{eff},\odot})^4$$

with $T_{\rm eff,\odot}$ the solar effective temperature, 5777 K. Calculate $D_{\rm mom}$ by providing \dot{M} in units of $10^{-6} M_{\odot}/yr$ and $v_{\rm inf}$ in units of km/s.

- Perform a linear fit to the (log-log) data, f(x) = ax + b, and overplot the linear regression.
- From these results, write down the WLR for Galactic supergiants (i.e, $\log D_{\rm mom} =$?) within your log-file
- Finally, add a *meaningful* title, axis-labels and legend.
- Plot the corresponding figure as a **ps file**, and inspect this file via **gv**.

Rename your files to 7_my_wlr.dat, 8_my_wlr.gpl, and 9_my_wlr.ps, respectively.

Finalizing the exercise

When you are finished with all your exercises, please check your log-file carefully, and provide a version which is easily readable and understandable. Of course, it is not forbidden to create a complete write-up via Latex or similar tools. Convert your log-file/write-up to pdf, and save it under the filename O_YourName_logfile.pdf.

Now, you should have (among others) the following files in your working directory:

```
0_YourName_logfile.pdf
1_Europe.pdf
2_wlr.dat
3_world.dem
4_world.ps
5_world_munich.ps
6_ostars.ps
7_my_wlr.dat
8_my_wlr.gpl
9_my_wlr.ps
```

and the directory Europe.

When you are satisfied with everything, *delete all other files and directories*, so that your working directory consists of these 10 files plus directory Europe only. If there are other files left or the files have the wrong names, this will have a significant impact on the grade for your exercise.

Tar and gzip your working directory to YourName_ex.tgz, and email this file to your supervisor.