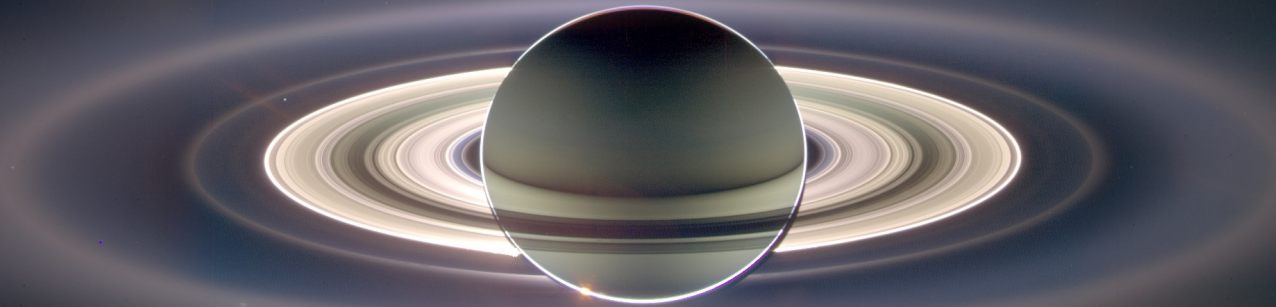


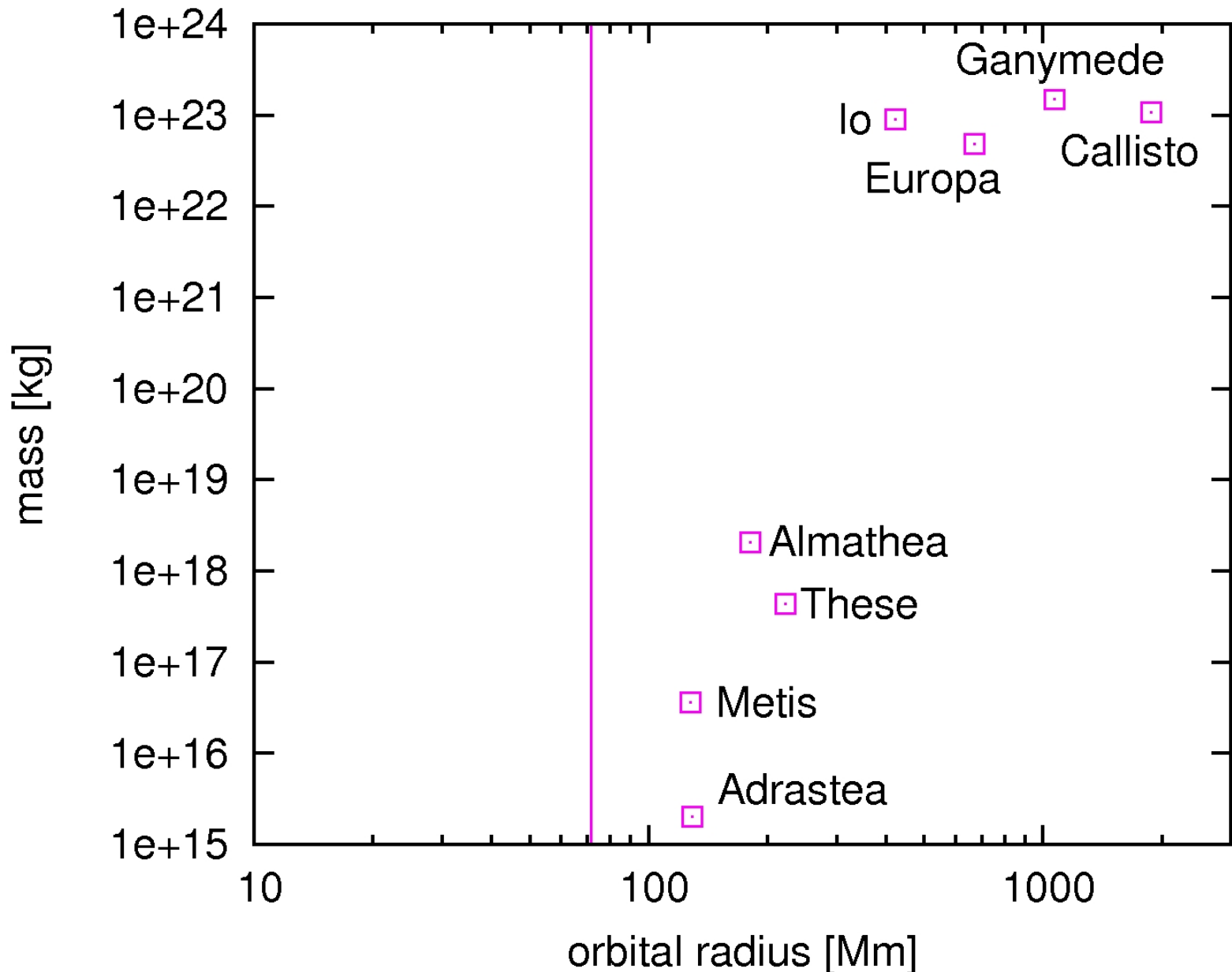
# FORMATION OF SATELLITES from a tidal disk



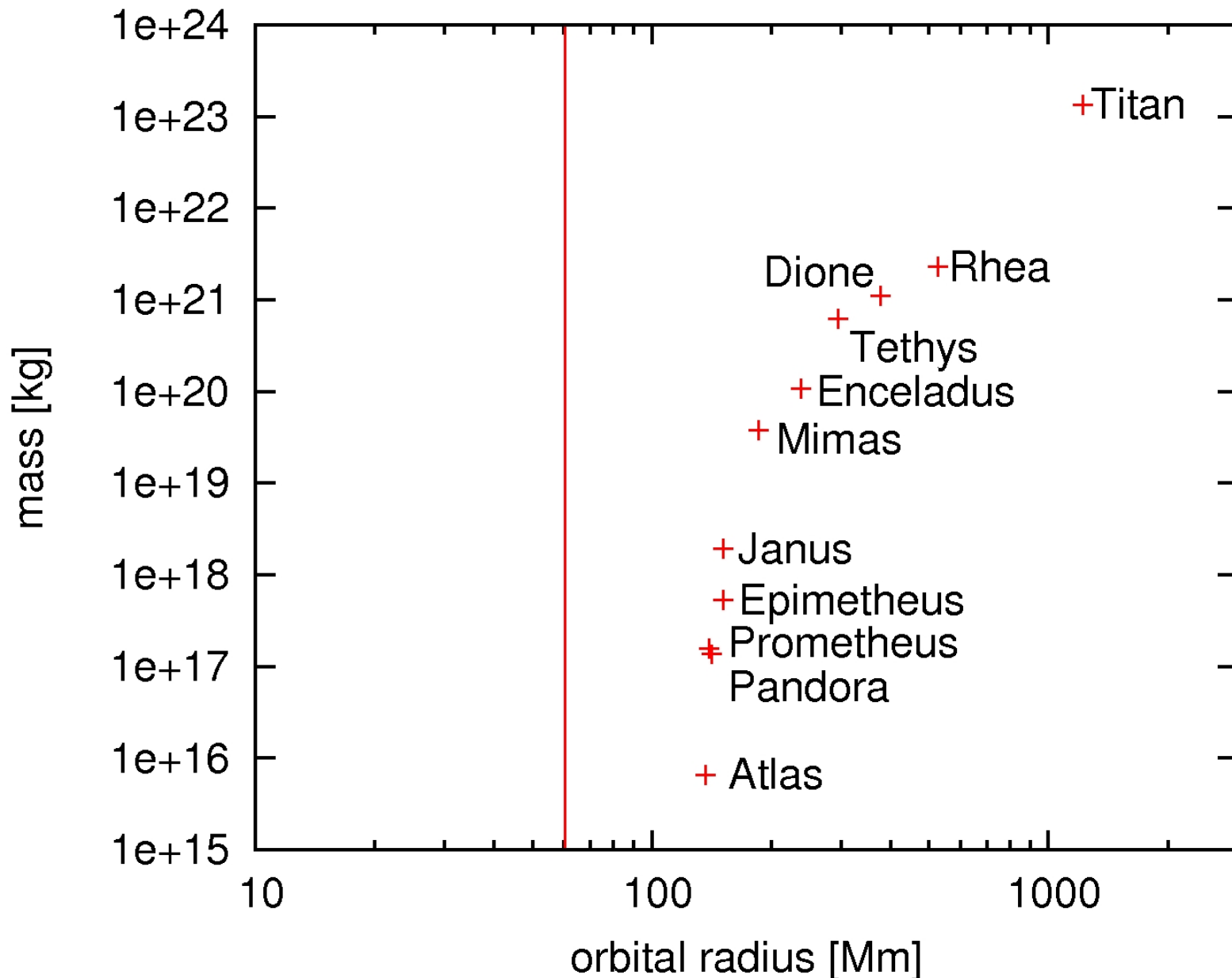
**Aurélien CRIDA**

& Sébastien CHARNOZ

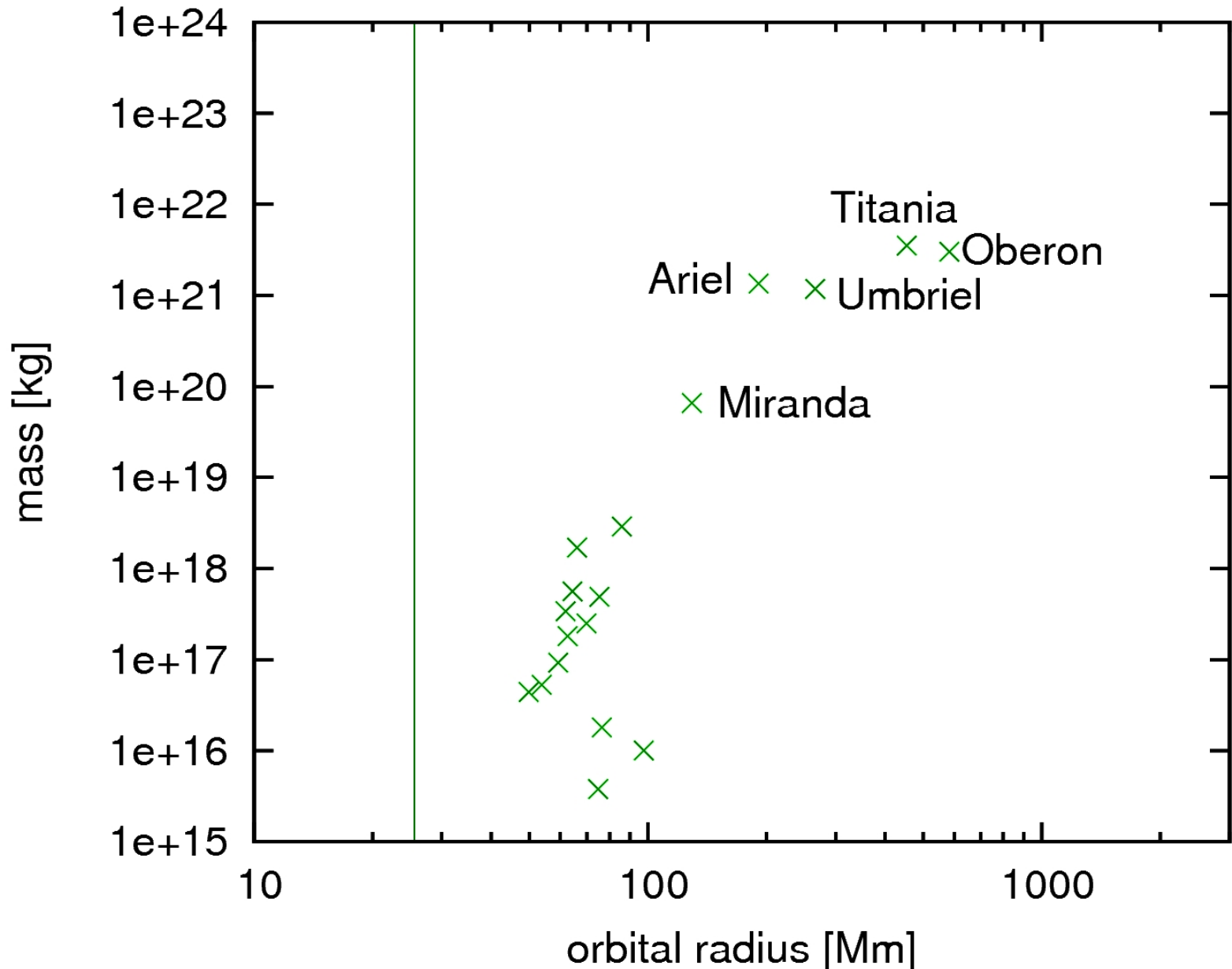
# JUPITER



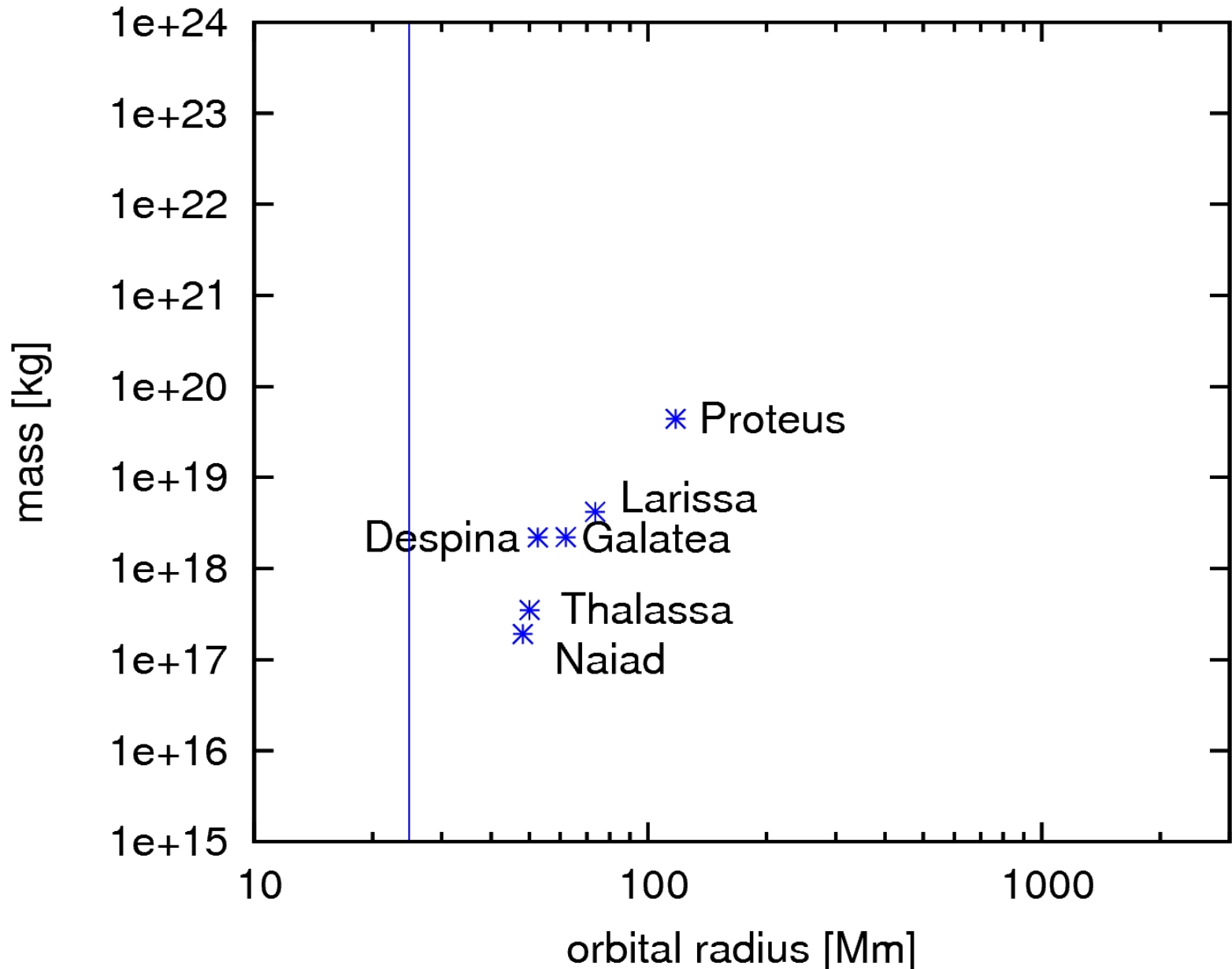
# SATURN



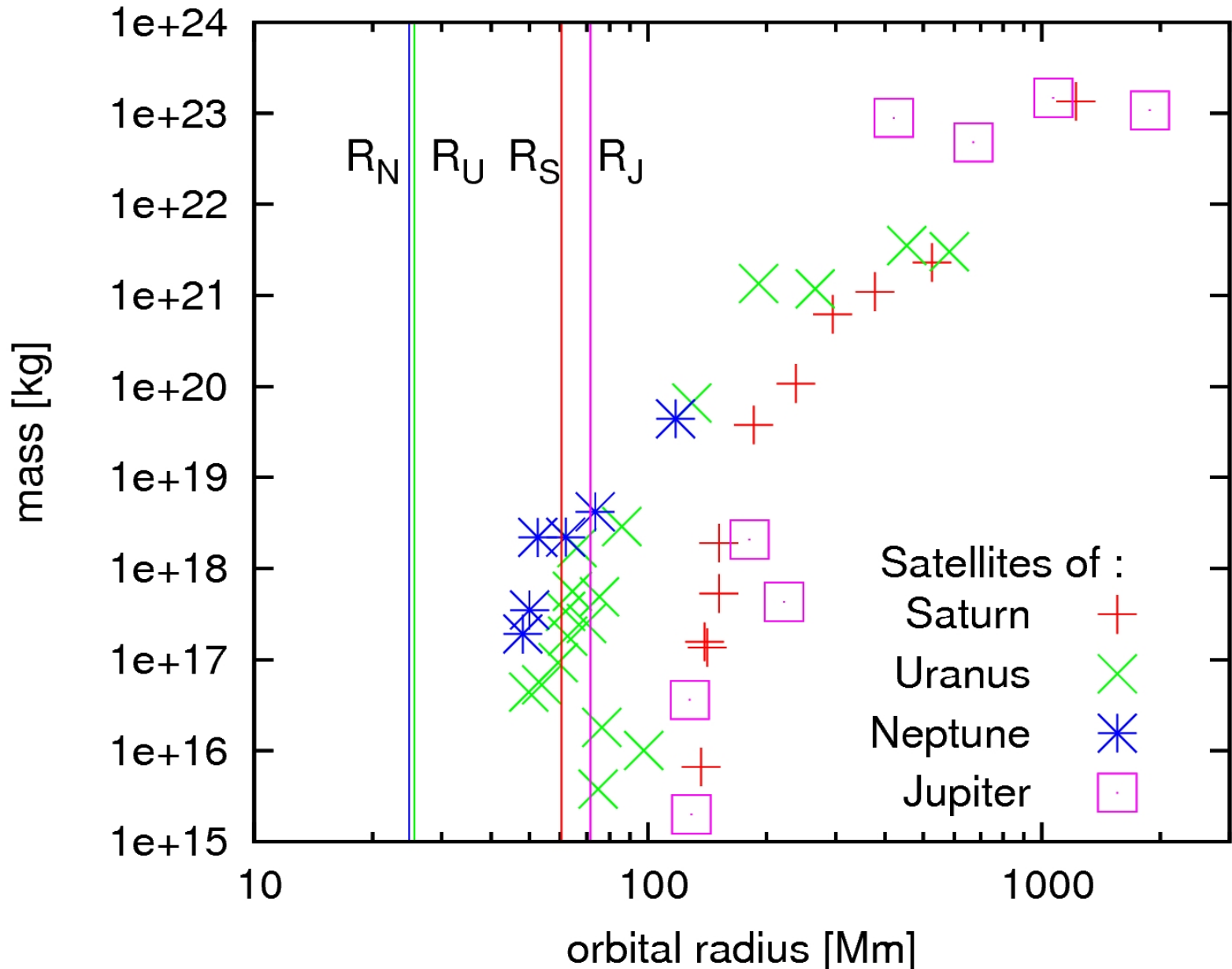
# URANUS



# NEPTUNE



# ALL GIANT PLANETS



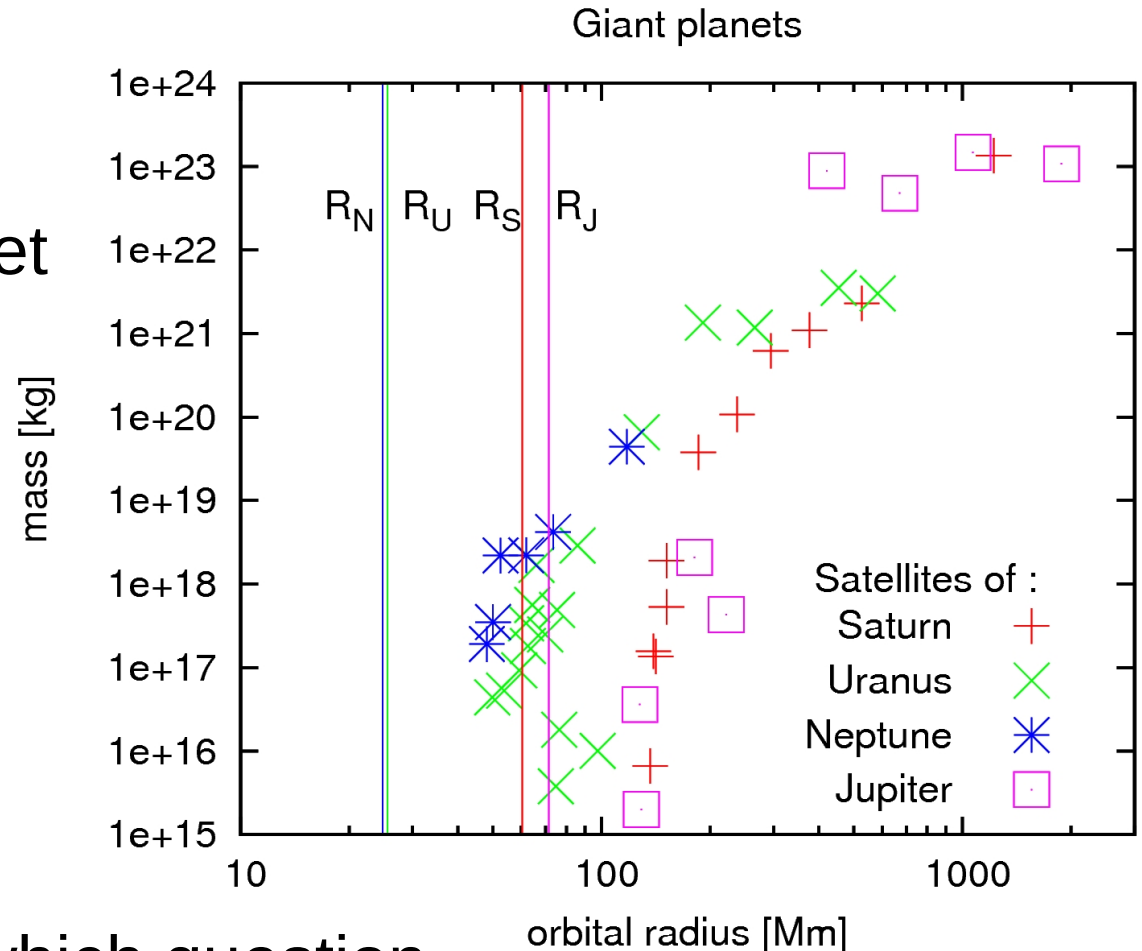
# INTRODUCTION

Distributions of giant planets' regular satellites :

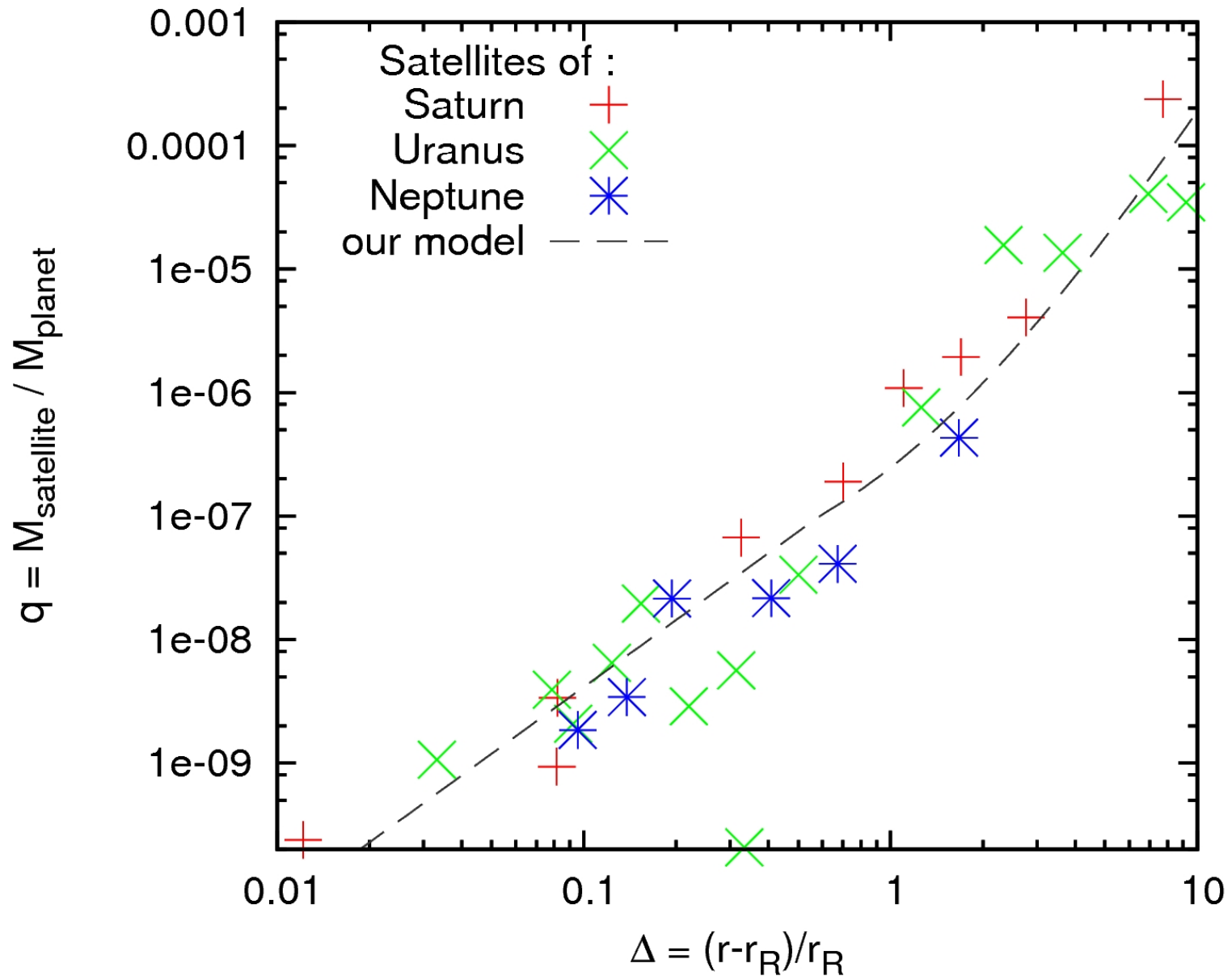
- don't reach the planet
- ranked by mass
- pile-up at a few planetary radii (small bodies)

Why ?

It's not a power law, which question the Circum-Planetary Disc model...



# CONCLUSION

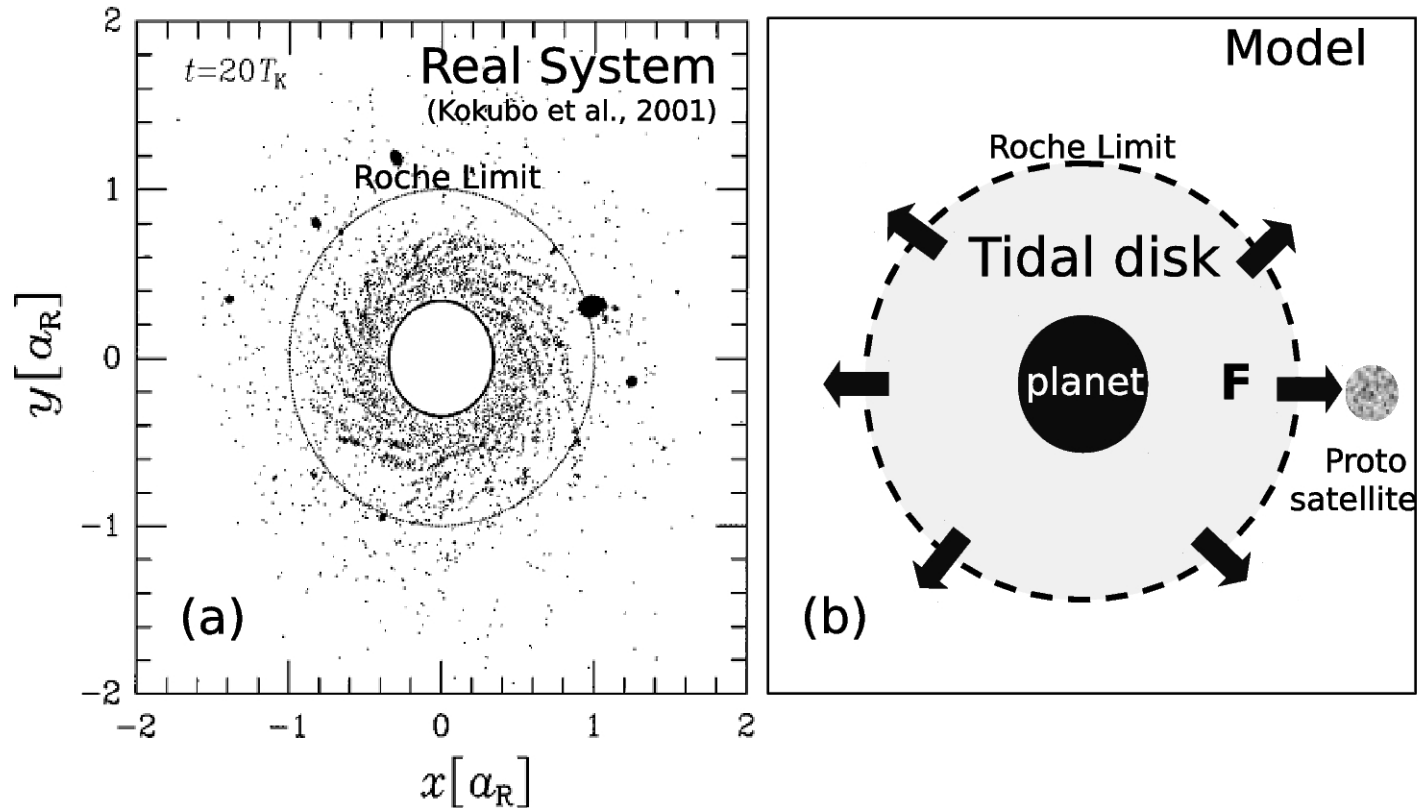




# Spreading of a tidal disk

1D model.

Inside the Roche radius  $r_R$ , there is a « tidal disk », that spreads with a mass flow  $F$  (assumed constant).



# Notations

Be  $T_R$  the orbital period at  $r_R$ , and

$\tau_{\text{disk}} = M_{\text{disk}} / FT_R$ , the normalized life-time of the disk.

The disk spreads with a viscous time  $t_v = r_R^2 / \nu$ .

Using Daisaka et al. (2001)'s prescription for  $\nu$ ,  
we find  $\tau_{\text{disk}} = t_v / T_R = 0.0425 D^{-2}$  where  $D = M_{\text{disk}} / M_p$ ,

and  $F = 23 D^3 M_p / T_R$ .

# Continuous regime

Say 1 satellite forms. Its mass is :  $M = F t$  (1)

It feels a torque from the tidal disk :  $\Gamma = \frac{8}{27} \left( \frac{M}{M_p} \right)^2 \Sigma r^4 \Omega^2 \Delta^{-3}$

where  $\Delta = (r - r_R) / r_R$  (Lin & Papaloizou 1979).

→ Migration rate :

where  $q = M / M_p$ .

$$\frac{d \Delta}{d t} = \frac{32}{27} q D T_R^{-1} \Delta^{-3} \quad (2)$$

Solution of (1) & (2) :

$$q = \left( \frac{\sqrt{3}}{2} \right)^3 \tau_{disk}^{-1/2} \Delta^2 \quad (3)$$

We call this the *continuous regime*.

# Continuous regime

This holds as long as the satellite captures immediately what comes through  $r_R$ .

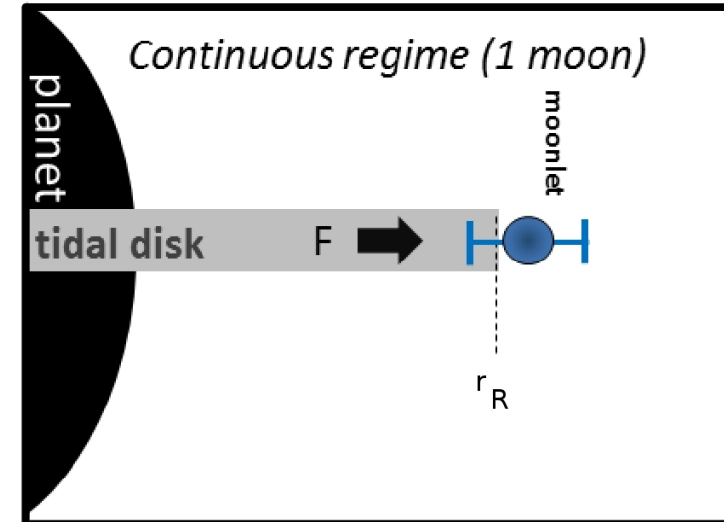
That is, as long as  $(r-r_R) < 2 r_{\text{Hill}}$ ,  
or  $\Delta < 2 (q/3)^{1/3}$ .

Input into Eq.(3), this gives a condition of validity for the continuous regime :

$$\Delta < \Delta_c = \sqrt{\frac{3}{\tau_{\text{disk}}}} = \sim 8.4 D$$

$$q < q_c = \frac{3^{5/2}}{2^3} \tau_{\text{disk}}^{-3/2} = \sim 222 D^3$$

Duration of the continuous regime:  $10 T_R$ .



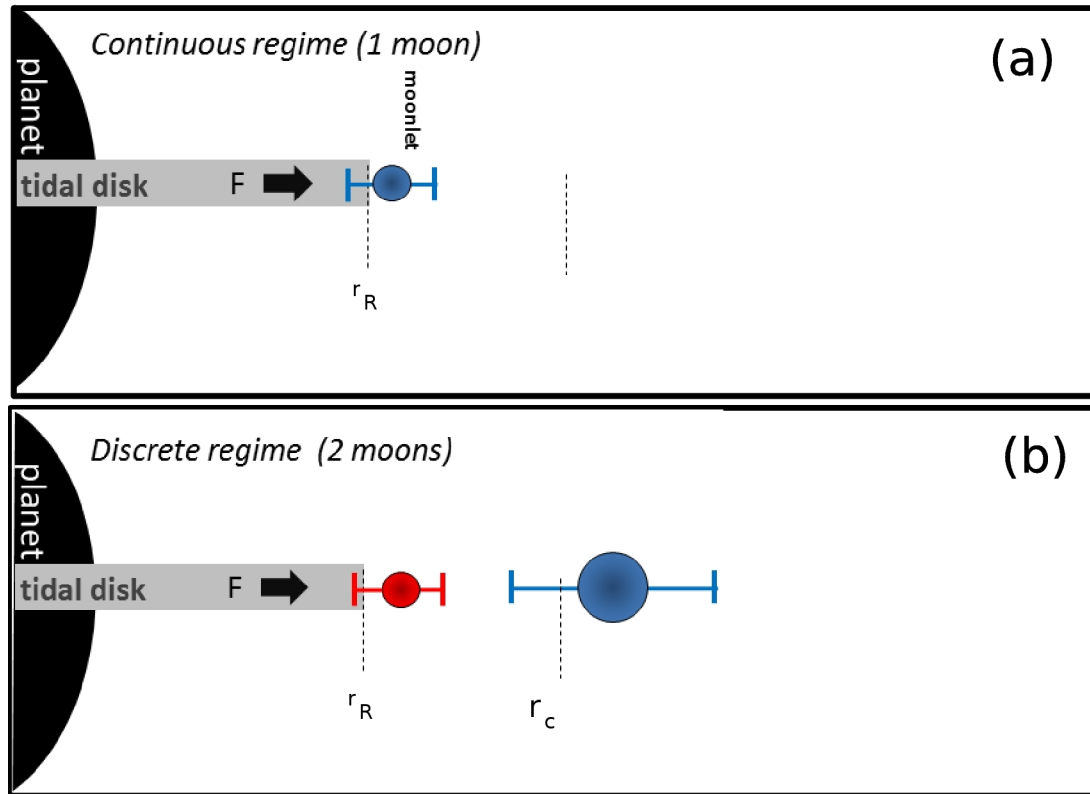
# Discrete regime

When the satellite is beyond  $\Delta_c$  (or  $q_c$ ), the material flowing through  $r_R$  forms a new satellite at  $r_R$ .

This new satellite is immediately accreted by the first one.

And so on...

The first satellite still grows as  $M=Ft$ , but by steps : *discrete regime*.



# Discrete regime

This holds as long as  $\Delta < \Delta_c + 2(q/3)^{1/3}$  .

It gives the condition :

$$\Delta < \Delta_d = 3.1 \Delta_c = \sim 26 D$$

$$q < q_d = 9.9 q_c = \sim 2200 D^3$$

The duration of the discrete regime is  $\sim 100 T_R$  .

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The duration of the discrete regime is  $\sim 100 T_R$  .

## Applications :

- 1) Earth's Moon forming disk :  $q_d = \sim$ mass of the Moon !
- 2) Charon never left the continuous regime.
- 3) Saturn's rings :  $q_d = \sim 10^{-18}$  .

# Pyramidal regime

Satellites of mass  $q_d$  are produced at  $\Delta_d$  every  $q_d / F$ .

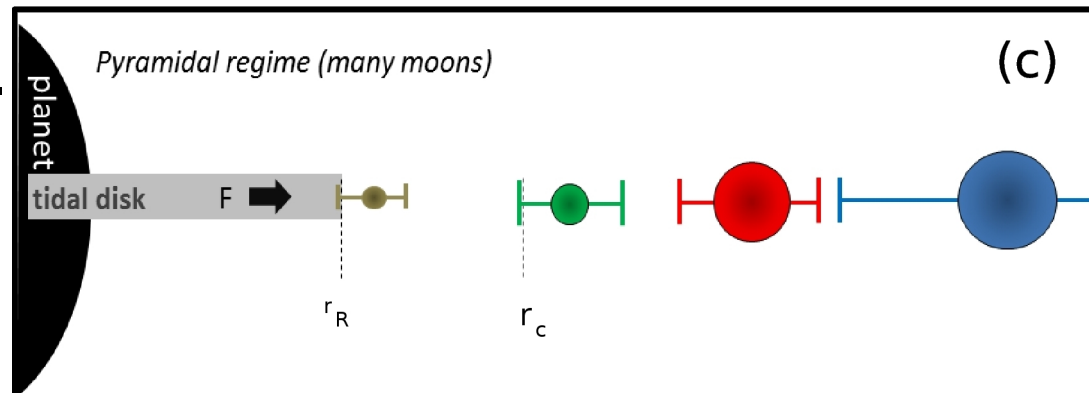
Then, many satellites of constant mass migrate outwards, at decreasing rates. They approach each other.

If their distance decreases below 2 mutual Hill radii, they merge.

This leads to the formation of satellites of masses  $2q_d$ , every  $2q_d / F$ . They migrate away and merge further...

And so on, hierarchically...

We call this *the pyramidal regime*.





# Pyramidal regime

- Using Eq.(2), we show that in the pyramidal regime, while the mass is doubled,  $\Delta$  is multiplied by  $2^{5/9}$ .

Thus,  $q \propto \Delta^{9/5}$ .

In addition, the number density of satellites should be proportionnal to  $1/\Delta$ , explaining the pile-up.

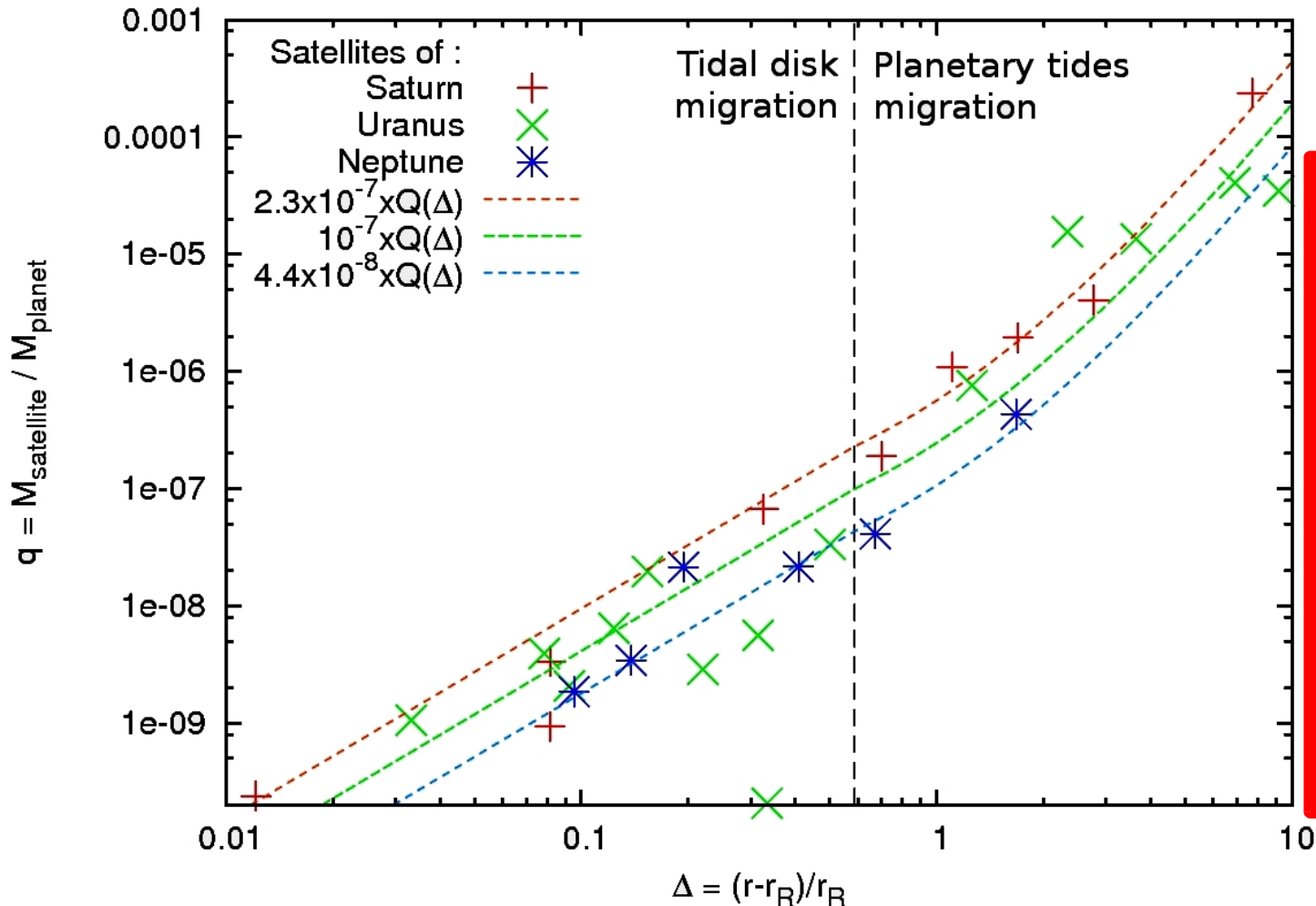
- Beyond the 2:1 Lindblad resonance with  $r_R$  ( $\Delta=0.58$ ), Eq.(2) doesn't apply. Migration is driven by planetary tides:

$$\frac{dr}{dt} = \frac{3 k_{2p} M \sqrt{G} R_p^5}{Q_p \sqrt{M_p} r^{11/2}} \quad (4)$$

Using Eq.(4), we find  $q \propto r^{3.9}$ .

# Pyramidal regime

The result spectacularly matches the distribution of the Saturnian, Uranian, and Neptunian systems !



I claim that Uranus and Neptune had massive rings, from which their regular satellites were born.

# Summary

## 1) Continuous regime:

1 moon grows

$$q \propto \Delta^2$$

until  $\Delta_c$  or  $q_c$ .

## 2) Discrete regime:

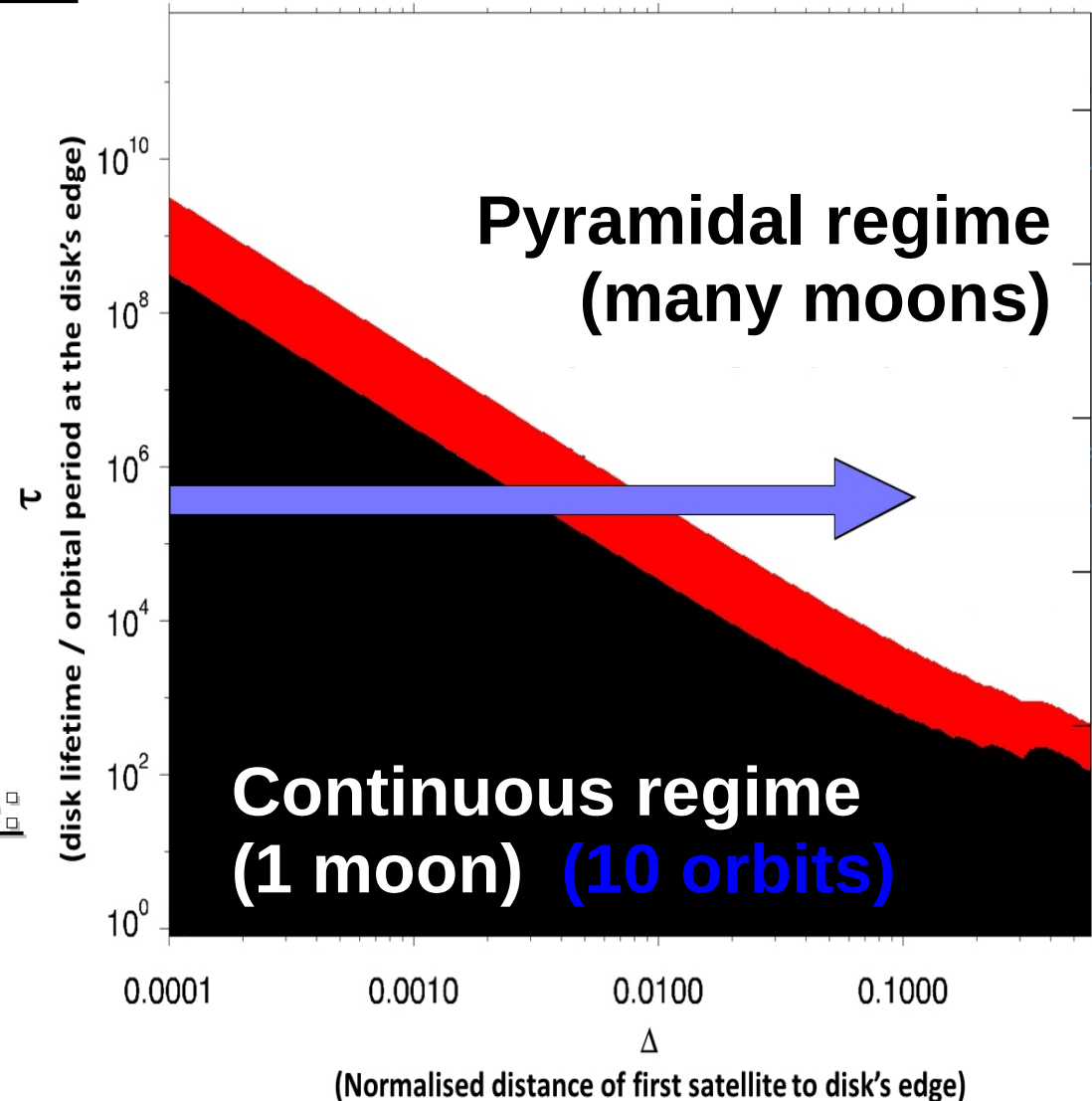
2 moons,  
growth by steps

until  $\Delta_d$  or  $q_d$ .

## 3) Pyramidal regime:

Many moons in the  
system.

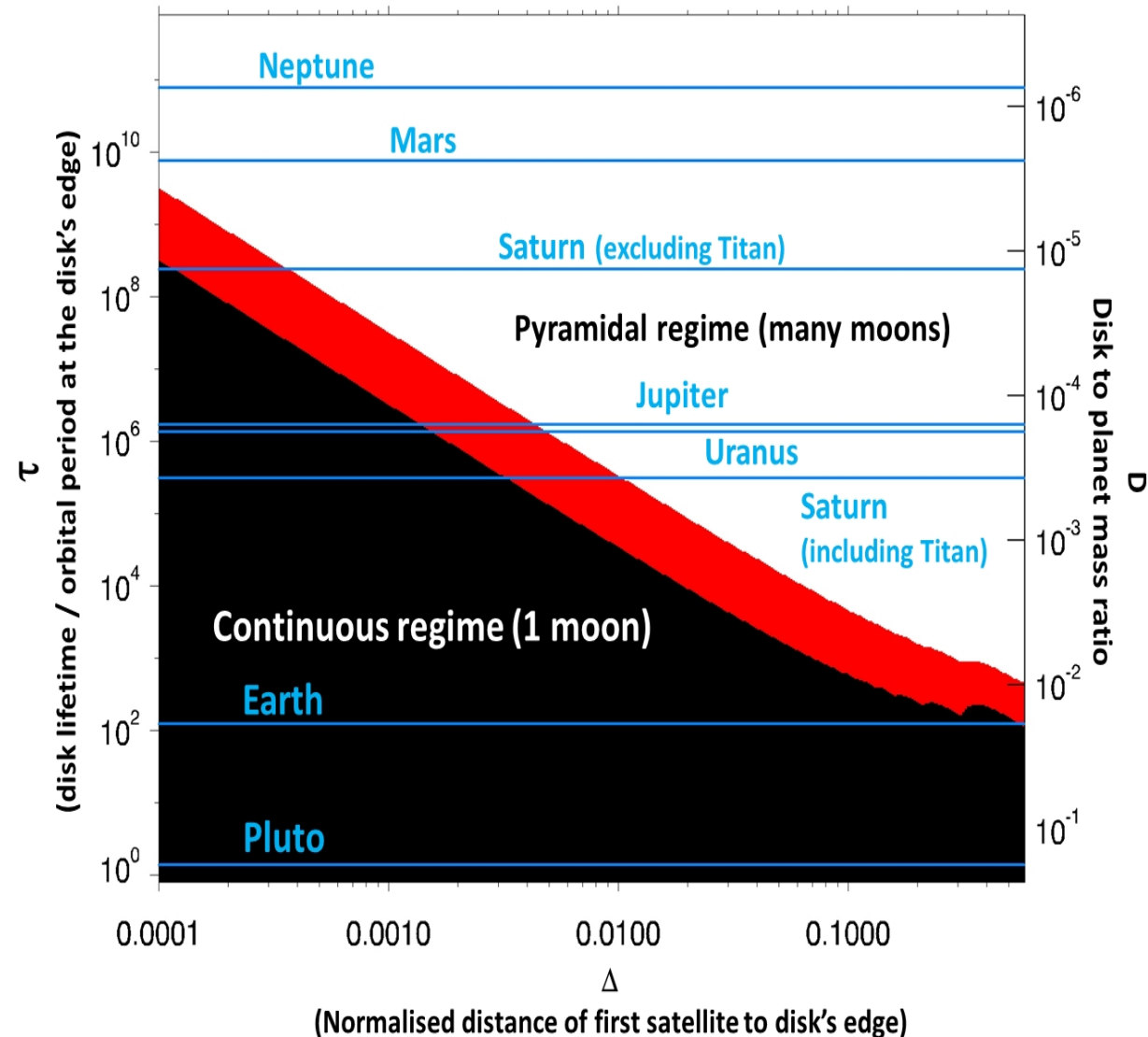
$$q \propto \Delta^{9/5} \text{ or } r^{3.8}.$$



# Summary

Take  $M_{\text{disk}} = 1.5 \times$   
the mass of the  
present satellite  
system.

Giant planets must  
be dominated by the  
pyramidal regime,  
while we expect the  
Earth and Pluto to  
have 1 large satellite.



# Conclusion & Discussion

- The spreading of a tidal disk beyond the Roche radius
- ✓ explains the mass-distance distribution of the regular satellites of the giant planets (observational signature of this process)
  - ✓ unifies terrestrial and giant planets in the same paradigm.
  - ✓ most Solar System regular satellites formed this way.
- ✗ Jupiter doesn't fit in this picture : probably formed in a circum-planetary disk (Canup & Ward 2002, 2006 ; Mosqueira & Estrada 2003a,b)
- Titan fits very well in this picture, though its « tidal age » is too large... Coincidence ?

Thanks !

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**Observatoire**  
de la CÔTE d'AZUR