

The envelope, torus or disk in the symbiotic binary

YY Her?

Bonn, December 12. - 13. 2006

Dr. Ladislav Hric

Astronomical Institute of the
Slovak Academy of Sciences
and

Dr. Rudolf Gális

University of P.J.Šafárik, Košice

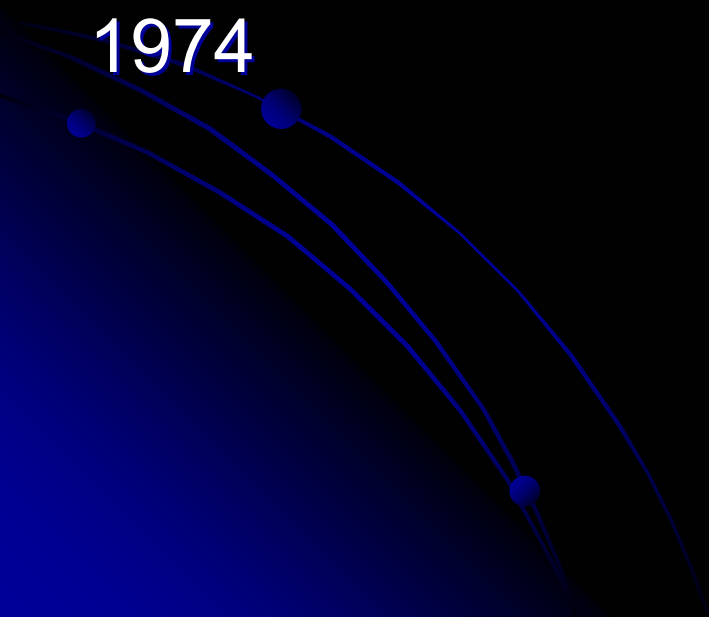
HARDY

Introduction

☀ YY Her belongs to the classical symbiotic binaries

💣 4 large outbursts 1914-1918, 1930-1933, 1981-1982, 1993-1996

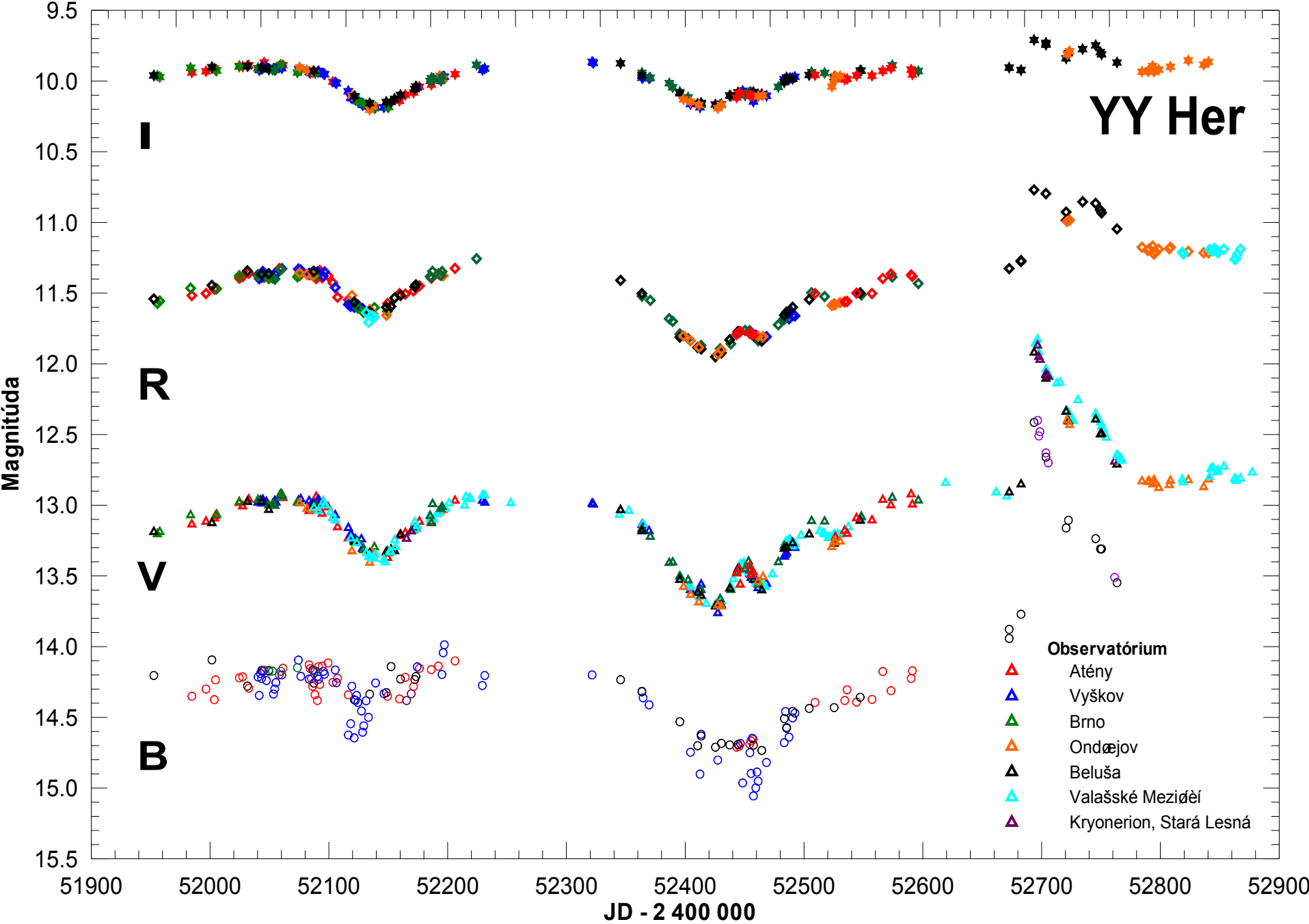
💣 6 small eruptions in 1890, 1903, 1942, 1954, 1965 a 1974



Our international photometric campaign

- Discovery of the secondary minimum
- Covering of the primary minimum
- Detection of the outburst activity
- Hric et al. (2001) new ephemeris


$$JD(I)_{\min} = 2450701,^{\text{d}}6 \pm 1,^{\text{d}}0 + 587,^{\text{d}}54 \pm 0,^{\text{d}}50 \times E$$

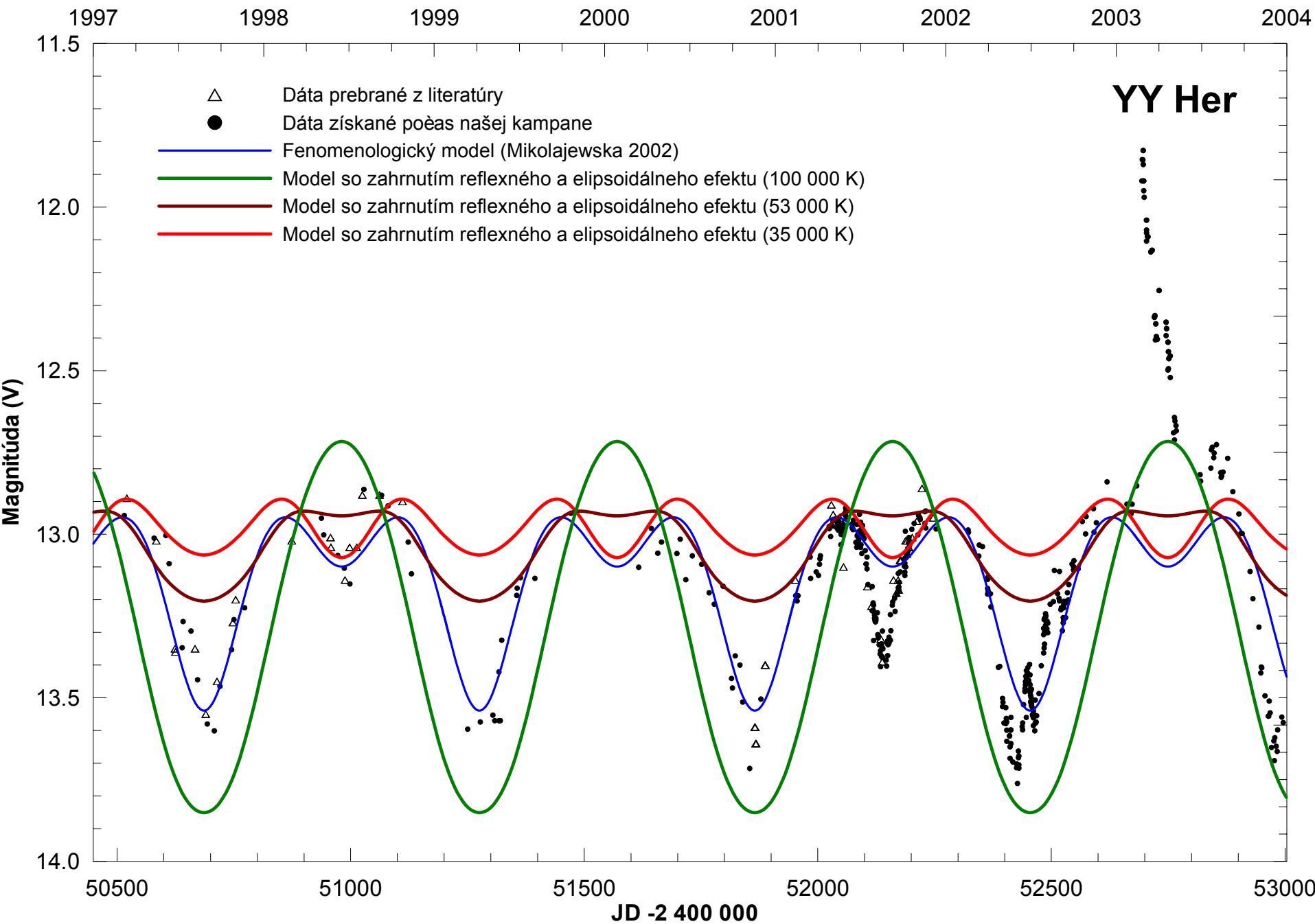


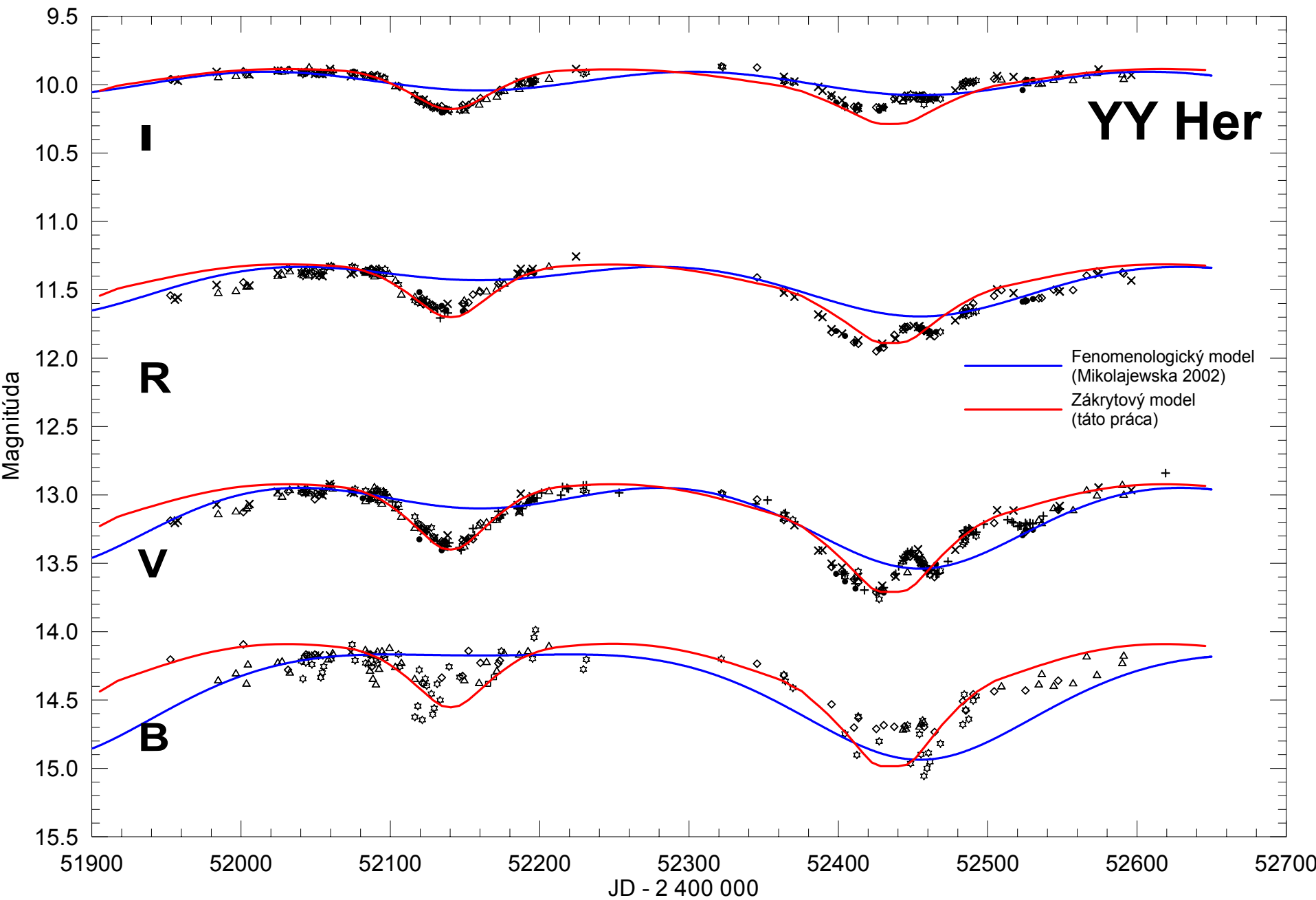
Recent research of YY Her

- Munari *et al.* (1997a) excluded eclipses as the cause of the light variability.
- Mikolajewska *et al.* (2002) explained the light variability of YY Her by combination of the ellipsoidal changes and sinusoidal variations of the nebular continuum and line emission.
- Formiggini & Leibowitz (2006) proposed to explain the secondary minimum by dark spots on the surface of the rotating red giant.

Our explanation

- **Eclipsing model**
 - $T_{\text{WD}} = 100\,000\text{ K}$, $T_{\text{c}} = 3\,500\text{ K}$,
 $q = M_{\text{c}} / M_{\text{WD}} = 2$
 - red giant near to its Roche lobe
 - white dwarf embedded in the envelope
with the temperature $T_{\text{en}} = 4\,000\text{ K}$
- 





Eclipsing model fits well the photometric behaviour of the system, but the question is: What nature and shape has the envelope around the white dwarf – is it envelope, torus or disk ?

The thickness of the envelope is 0.27 times of its radius, i. e. $H_{\text{en}} = 27 R_{\odot}$
very probably disk



Thank you for your attention

Bonn, December 12. - 13. 2006